How to prepare the interface between SIESTA and WANNIER90





Javier Junquera



Important bibliography:

For a review on Maximally Localized Wannier functions:

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Maximally localized Wannier functions: Theory and applications

Nicola Marzari

Theory and Simulation of Materials (THEOS), École Polytechnique Fédérale de Lausanne, Station 12, 1015 Lausanne, Switzerland

Arash A. Mostofi

Departments of Materials and Physics, and the Thomas Young Centre for Theory and Simulation of Materials, Imperial College London, London SW7 2AZ, United Kingdom

Jonathan R. Yates

Department of Materials, University of Oxford, Parks Road, Oxford OX1 3PH, United Kingdom

Ivo Souza

Centro de Física de Materiales (CSIC) and DIPC, Universidad del País Vasco, 20018 San Sebastián, Spain and Ikerbasque Foundation, 48011 Bilbao, Spain

David Vanderbilt

Department of Physics and Astronomy, Rutgers University, Piscataway, New Jersey 08854-8019, USA

Last features implemented in the WANNIER90 code

http://www.wannier.org

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Wannier90 as a community code: new features and applications

Giovanni Pizzi^{1,29,30}, Valerio Vitale^{2,3,29}, Ryotaro Arita^{4,5}, Stefan Blügel⁶, Frank Freimuth⁶, Guillaume Géranton⁶, Marco Gibertini^{1,7}, Dominik Gresch⁸, Charles Johnson⁹, Takashi Koretsune^{10,11}, Julen Ibañez-Azpiroz¹², Hyungjun Lee^{13,14}, Jae-Mo Lihm¹⁵, Daniel Marchand¹⁶, Antimo Marrazzo¹⁰, Yuriy Mokrousov^{6,17}, Jamal I Mustafa¹⁸, Yoshiro Nohara¹⁹, Yusuke Nomura⁴, Lorenzo Paulatto²⁰, Samuel Poncé²¹, Thomas Ponweiser²², Junfeng Qiao²³, Florian Thöle²⁴, Stepan S Tsirkin^{12,25}, Małgorzata Wierzbowska²⁶, Nicola Marzari^{1,29}, David Vanderbilt^{27,29}, Ivo Souza^{12,28,29}, Arash A Mostofi^{3,29}, and Jonathan R Yates^{21,29},

Important bibliography:

The user guide and tutorial of the WANNIER90 code

wannier90: Tutorial

Version 3.1

Freely available from: http://www.wannier.org

After running SIESTA and compute the PDOS, we can analyze the character of the different bands

Which atoms contribute more to the bands at a particular energy window



We can analyze the character of the different bands

Which atoms contribute more to the bands at a particular energy window Zoom around the top of the valence bands and bottom of conduction bands



Bottom of conduction bands: mostly Ti character

Top of valence bands: mostly O character

We can project on particular atomic orbitals within an atom to further define the character.

Flowchart to compute Maximally Localized Wannier functions from the input provided by SIESTA

1. Prepare compatible input files for SIESTA and WANNIER90

2. Run WANNIER90 in pre-processing format

3. Run SIESTA and generate the required files to feed WANNIER90

4. Change the extension to the SystemLabel.eigW file

5. Run WANNIER90 to minimize the localization functional, plot the Wanniers, interpolate bands, etc as a post-processing tool

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Choose the Bloch states that will be used to compute the Wannier functions



In this particular example, we are interested in the wannierization of:

Bottom of conduction bands: mostly Ti character

Top of valence bands: mostly O character

That means: - 3 O p bands (p_x, p_y, p_z) × 3 O atoms - 3 Ti t_{2g} bands (d_{xy}, d_{yz}, d_{xz})

12 bands to wannierize

Prepare compatible input files for SIESTA and WANNIER90: How to select the number of bands to wannierized

seedname.win file (input of WANNIER90)

<pre>num_bands = 12 num_wann = 12 ! ! For every k-point, the second seco</pre>	<pre>! Number of bands in the first-principles calculation ! used to form the overlap matrices ! seedname.mmn file ! Default: num_bands = num_wann ! Number of Wannier functions to be found ! There must be as many projectors in the ! projections bloch as num_wann functions here. ! If num_wann < num_bands, then the disentaglement ! procedure is activated ! In this example, we want to Wannierize the ! top of the valence bands (O 2p in character) ! and the bottom of the conduction bands (Ti t2g), ! so it amounts to ! - O p bands (3 bands * 3 O atoms) ! - Ti t2g bands (3 bands: dxy, dyz, dxz) ! Twelve bands in total me bands in Siesta are ordered from lowest to highest</pre>	
<pre>! energy ! The character of the ! In order, for the cas ! ! Bands to be excluded</pre>	bands are known after computation of the PDOS. se of SrTiO3:	
! ! Band 1 (lowest band) ! Bands 2, 3, and 4: ! Band 5: ! Bands 6, 7, and 8: ! Band 9, 10, and 11:	First Scharacter Ti-3s character (px, py, and pz) Sr-4s character O-2s character (1 band * 3 O atoms in the unit cell) Sr-4p character (px, py, and pz)	
! Bands to be wannieri:	zed	
! Bands 12-20: ! Bands 21, 22, and 23	O-2p character (3 bands * 3 O atoms in the unit cell) Ti-t2g characer (dxy, dyz, dxz)	
<pre>! In exclude_bands, we list the bands to excluded from the calculation ! Here we exclude all bands except the 02p and Ti t2g bands. ! That means that we exclude: ! - Ti 3s band (1 band) ! - Ti 3p bands (3 bands) ! - Sr 4s bands (1 band) ! - Sr 4p bands (3 bands) ! - 0 2s bands (1 band * 3 atoms = 3 bands) ! This amounts to the first eleven bands for every k-point.</pre>		

seedname.fdf file (input of SIESTA)

Siesta2Wannier90.WriteMmn Siesta2Wannier90.WriteAmn Siesta2Wannier90.WriteEig Siesta2Wannier90.WriteUnk		ເe. ເe. ເe.
Siesta2Wannier90.NumberOfBands	23	
	00	•

Siesta2Wannier90.UnkGrid1	20
Siesta2Wannier90.UnkGrid2	20
Siesta2Wannier90.UnkGrid3	20

Select the same number of bands here as required to include all the bands.

Internally, SIESTA discards those listed in exclude_bands and produce the overlap matrices only for the bands to be wannierized

Prepare compatible input files for SIESTA and WANNIER90: The atomic geometry

The unit cells lattice vectors and atomic positions must be the same in the WANNIER90 input and in the SIESTA input used to generate the projections

seedname.win file (input of WANNIER90)

seedname.fdf file (input of SIESTA)

begin unit_cell_cart	LatticeConstant 3.874 Ang
Ang	%block LatticeVectors
3.874 0.000 0.000	1.000 0.000 0.000
0.000 3.874 0.000	0.000 1.000 0.000
0.000 0.000 3.874	0.000 0.000 1.000
end unit_cell_cart	%endblock LatticeVectors
begin atoms_frac Sr 0.0 0.0 0.0 Ti 0.5 0.5 0.5 0 0.5 0.5 0.0 0 0.5 0.0 0.5 0 0.0 0.5 0.5 end atoms frac	AtomicCoordinatesFormat Fractional %block AtomicCoordinatesAndAtomicSpecies 0.00000000 0.00000000 0.00000000 1 87.62 Sr 0.50000000 0.50000000 0.50000000 2 47.867 Ti 0.50000000 0.50000000 0.00000000 3 15.9994 0 0.50000000 0.50000000 0.50000000 3 15.9994 0 0.00000000 0.50000000 0.50000000 3 15.9994 0 %endblock AtomicCoordinatesAndAtomicSpecies

The unit cell lattice vectors must be the same in symmetry and length of the cell vectors

The coordinates of the atoms must be the same, as well as the order in which they are entered

Prepare compatible input files for SIESTA and WANNIER90: The WANNIER90 input file: k-points block

SIESTA must compute the overlap of the periodic part of the wave functions at neighbour k-points

$$M_{mn}^{(k,b)} = \langle u_{m\vec{k}} | u_{n\vec{k}+\vec{b}} \rangle$$

 \vec{b} is the vector which connect a given \vec{k} with its neighbours

seedname.win file (input of WANNIER90) to indicate which \vec{k} points will be considered

mp_grid : 4 4 4

begin kpoints

0.00000000	0.00000000	0.0000000
0.00000000	0.00000000	0.25000000
0.00000000	0.00000000	0.5000000
0.00000000	0.0000000	0.75000000
0.00000000	0.25000000	0.0000000
0.00000000	0.25000000	0.25000000
0.00000000	0.25000000	0.5000000
0.00000000	0.25000000	0.75000000
0.00000000	0.50000000	0.0000000
0.00000000	0.50000000	0.25000000
0.00000000	0.50000000	0.5000000
0.00000000	0.50000000	0.75000000
0.00000000	0.75000000	0.0000000
0.00000000	0.75000000	0.25000000
0.00000000	0.75000000	0.5000000
0.00000000	0.75000000	0.75000000
• • •		
0.75000000	0.75000000	0.75000000
nd kpoints		

Dimensions of the regular Monkhorst-Pack \vec{k} -point grid mesh

Each line gives the coordinate of a \overline{k} -point in relative units, i.e. in fractional units with respect to the primitive reciprocal lattice vector

This list of \vec{k} -points is independent of the Monkhorst-Pack mesh used to achieve the self-consistency

Prepare compatible input files for SIESTA and WANNIER90: The WANNIER90 input file: k-points block

SIESTA must compute the overlap of the periodic part of the wave functions at neighbour k-points

 $M_{mn}^{(\vec{k},\vec{b})} = \langle u_{m\vec{k}} | u_{n\vec{k}+\vec{b}} \rangle$

seedname.win file (input of WANNIER90) to indicate which \vec{k} points will be considered

mp_grid : 4 4 4

begin kpoints

0.00000000	0.00000000	0.00000000
0.0000000	0.0000000	0.25000000
0.00000000	0.00000000	0.5000000
0.0000000	0.0000000	0.75000000
0.00000000	0.25000000	0.0000000
0.0000000	0.25000000	0.25000000
0.0000000	0.25000000	0.5000000
0.00000000	0.25000000	0.75000000
0.0000000	0.50000000	0.0000000
0.0000000	0.50000000	0.25000000
0.00000000	0.50000000	0.5000000
0.00000000	0.50000000	0.75000000
0.0000000	0.75000000	0.0000000
0.00000000	0.75000000	0.25000000
0.0000000	0.75000000	0.5000000
0.0000000	0.75000000	0.75000000
• • •		
0.75000000	0.75000000	0.75000000
nd kpoints		

 \vec{b} is the vector which connect a given \vec{k} with its neighbours

seedname.fdf file (input of SIESTA)

Siesta2Wannier90.WriteMmn	.true.
Siesta2Wannier90.WriteAmn	.true.
Siesta2Wannier90.WriteUnk	.true.
Siesta2Wannier90.NumberOfBands	23
Siesta2Wannier90.UnkGrid1 Siesta2Wannier90.UnkGrid2 Siesta2Wannier90.UnkGrid3	20 20 20

The information is passed through the seename.nnkp file, generated by WANNIER90 after running in pre-processing mode

Prepare compatible input files for SIESTA and WANNIER90: The WANNIER90 input file: projections

SIESTA must compute the overlap of the Bloch states $|\psi_{m\vec{k}}
angle$ onto trial localised orbitals $|g_n
angle$

 $A_{mn}^{(\vec{k})} = \langle \psi_{m\vec{k}} | g_n \rangle$

Each projection is associated with a site and an angular momentum state defining the projection function

seedname.win file (input of WANNIER90) to indicate the projection functions

In our example, we are going to project over: the *p*-states of all the O and the t_{2p} orbitals of Ti

<pre>guiding_centres = T</pre>	! !	The projection centres are used as the guiding centres in the Wannierisation routines
begin projections O:p Ti:1=2,mr=2,3,5 end projections	! ! !	The projections block defines a set of localised functions used to generate an initial guess for the unitary transformations.

Take a look to the WANNIER90 User's Guide (pages 33-38) to check the syntaxis and short-cuts

Prepare compatible input files for SIESTA and WANNIER90: The WANNIER90 input file: projections

SIESTA must compute the overlap of the Bloch states $|\psi_{m\vec{k}}
angle$ onto trial localised orbitals $|g_n
angle$

 $A_{mn}^{(\vec{k})} = \langle \psi_{m\vec{k}} | g_n \rangle$

Each projection is associated with a site and an angular momentum state defining the projection function

seedname.win file (input of WANNIER90) to indicate the projection functions

guiding_centres = T

begin projections 0:p Ti:l=2,mr=2,3,5 end projections seedname.fdf file
(input of SIESTA)

Siesta2Wannier90.WriteMmn	.true.
Siesta2Wannier90.WriteAmn	.true.
Siesta2Wannier90.WriteEig Siesta2Wannier90.WriteUnk	.true. .true.
Siesta2Wannier90.NumberOfBands	23
Siesta2Wannier90.UnkGrid1 Siesta2Wannier90.UnkGrid2 Siesta2Wannier90.UnkGrid3	20 20 20

The information is passed through the seename.nnkp file, generated by WANNIER90 after running in pre-processing mode

Flowchart to compute the Maximally Localized Wannier functions from the input provided by siesta

1. Prepare compatible input files for SIESTA and WANNIER90

2. Run WANNIER90 in pre-processing format

\$ wannier90.x -pp SrTiO3

This generates a seedname.nnkp file that is digested by SIESTA

Flowchart to compute the Maximally Localized Wannier functions from the input provided by siesta

1. Prepare compatible input files for SIESTA and WANNIER90

2. Run WANNIER90 in pre-processing format

\$ wannier90.x -pp SrTiO3

3. Run SIESTA and generate the required files to feed WANNIER90

\$ siesta < SrTiO3.fdf > SrTiO3.out

Succesful output of SIESTA

siesta2wannier90: Reading the SrTiO3.nnkp file read_nnkp: Checking info from the SrTiO3.nnkp file read_nnkp: Reading data about real lattice read_nnkp: - Real lattice is ok read_nnkp: Reading data about reciprocal lattice - Reciprocal lattice is ok read_nnkp: read_nnkp: Reading data about k-points read_nnkp: Reading data about projection centers read_nnkp: Reading data about k-point neighbours read_nnkp: Reading data about excluded bands noccbands: Total number of electrons 40 noccbands: Number of occupied bands 20 number_bands_wannier: Number of bands for wannierization number_bands_wannier: before excluding bands 23 0 = Number of bands for wannierization after excluding bands: 12 0 Bands to be wannierized: 12 13 14 15 16 17 18 19 20 21 22 23 compute_pw_matrix: Computing the matrix elements of a plane wave mmn: Overlap matrices between periodic part of wavefunctions mmn: written in SrTiO3.mmn file amn: Overlap matrices between trial projection functions and wavefunctions amn: written in SrTiO3.amn file eig: Eigenvalues of the Hamiltonian eig: written in SrTiO3.eigW file siesta2wannier90: All the information dumped in the corresponding files siesta2wannier90: End of the interface between Siesta and Wannier90

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3. Run SIESTA and generate the required files to feed WANNIER90

\$ siesta < SrTiO3.fdf > SrTiO3.out

4. Change the extension to the SystemLabel.eigW file

\$ mv SrTiO3.eigW SrTiO3.eig
overwrite SrTiO3.eig? (y/n [n]) y

Flowchart to compute the Maximally Localized Wannier functions from the input provided by siesta

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overwrite SrTiO3.eig? (y/n [n]) y

5. Run WANNIER90 to minimize the localization functional, plot the Wanniers, interpolate bands, etc as a post-processing tool

\$ wannier90.x SrTiO3

Successful output of WANNIER90

Cvcle: 100			
WF centre and spread	1 (1.937000, 1.937000, -0.000000)	1.14103775	
WF centre and spread	2 (1.937000, 1.937000, -0.000002)	1.01254313	
WF centre and spread	3 (1.937000, 1.937000, 0.000000)	1.01253811	
WF centre and spread	4 (1.936999, 0.000001, 1.937000)	1.01254138	9 WF centered
WF centre and spread	5 (1.937001, 0.000000, 1.937000)	1.01254288	on the three O
WF centre and spread	6 (1.937001, -0.000001, 1.937001)	1.14104145	
WF centre and spread	7 (0.000001, 1.937000, 1.937000)	1.01254154	
WF centre and spread	8 (-0.000001, 1.937001, 1.937000)	1.14104851	
WF centre and spread	9 (-0.000001, 1.937001, 1.937000)	1.01253866	
WF centre and spread	10 (1.937000, 1.937001, 1.937000)	0.90476826	
WF centre and spread	11 (1.937000, 1.937000, 1.936999)	0.90476759	3 WF centered
WF centre and spread	12 (1.937000, 1.937000, 1.937000)	0.90477667	on Ti
Sum of centres and sp	reads(17.433000, 17.433003, 17.432997) 12.21268	593
	-	-	
100 -0 178F-14 0 (000001797 12 2126859319 1 04	CONV	

Check that the minimization has converged

How to plot the Wannier functions

First of all, SIESTA has to write the periodic part of the Bloch functions in a 3D grid. The number of points in th grid along the three lattice vectors are given by

seedname.fdf file

(input of SIESTA)

Siesta2Wannier90.UnkGrid1	20
Siesta2Wannier90.UnkGrid2	20
Siesta2Wannier90.UnkGrid3	20

This produces many files with the name UNKXXXXX.Y where

- XXXXX is the number of the k-point, from 1 to the number of points included in seedname.win file
- Y refers to the spin compoment (1 or 2)

seedname.win file
(input of WANNIER90)

!To plot the WF	
wannier_plot = T	! If true, then the code will write out the
	! Wannier functions in a super-cell whose size is
	! and in a format specified by wannier_plot_format
<pre>wannier_plot_supercell = 3</pre>	
<pre>wannier_plot_list = 1,10</pre>	! A list of WF to plot.
	! The WF numbered as per the seedname.wout
	! file after the minimisation of the spread.
restart = plot	

\$ wannier90.x SrTiO3

How to plot the Wannier functions

WANNIER90 produces files with the name: SrTiO3_00002.xsf, SrTiO3_00010.xsf that can be directly plotted with XCRYSDEN

Once XCrySDen starts, click on Tools \rightarrow Data Grid

Click on OK Then, select: Degree of triCubic Spline: 3 Click on Render+/- isovalue Select the desired isovalue (in this example 0.1) Submit

SrTiO3_00002.xsf



SrTiO3_00010.xsf



How to plot the interpolated band structure obtained from the subspace spanned by the Wannier function

Seedname.win file (input of WANNIER90)

begin kpoint_path	! Defines the path in k-space along
G 0.000 0.000 0.000 X 0.500 0.000 0.000	! which to calculate the
X 0.500 0.000 0.000 M 0.500 0.500 0.000	! bandstructure. Each line gives the
M 0.500 0.500 0.000 G 0.000 0.000 0.000	! start and end point (with labels)
G 0.000 0.000 0.000 R 0.500 0.500 0.500	! for a section of the path.
R 0.500 0.500 0.500 X 0.500 0.000 0.000	! Values are in fractional
end kpoint_path	! coordinates with respect to the
	! primitive reciprocal lattice
	! vectors.
bands_plot =T	

\$ wannier90.x SrTiO3

This will produce three files that can be directly plotted by gnuplot SrTiO3_band.gnu SrTiO3_band.kpt SrTiO3_band.dat (in the last version of gnuplot, it is very likely that you have to edit SrTiO3_band.gnu and change the first line to set style data dots

How to plot the interpolated band structure obtained from the subspace spanned by the Wannier function

\$ gnuplot \$ load "SrTiO3_band.gnu"





How to plot the interpolated band structure obtained from the subspace spanned by the Wannier function

From SIESTA -10 -20 Energy (eV) -30 -40 -50 -60 -70 -10 -5 X R Х Μ Г Г DOS (eV^{-1})

Interpolated from WANNIER90



Only the top of the valence band and the bottom of the conduction band

Funding

SPANISH INITIATIVE FOR ELECTRONIC SIMULATIONS WITH THOUSANDS OF ATOMS: CÓDIGO ABIERTO CON GARANTÍA Y SOPORTE PROFESIONAL: SIESTA-PRO

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